Short communication

Carbon isotopic fractionation of carbon dioxide between gas and hydrate phases



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ABSTRACT. During the formation of gas hydrate crystals, stable isotope fractionation of the guest molecules occurs. For methane and ethane, hydrogen stable isotope fractionation has been reported in the range of a few per mil, but for carbon stable isotope fractionation, there is no difference in δ^{13} C, or the difference is below the detection limit. As for carbon dioxide in natural gas hydrates, little is known about stable isotope fractionation of carbon dioxide. In this study, we report the stable isotope fractionation during formation of synthetic carbon dioxide hydrates under various temperature conditions. Fine ice powder and pure carbon dioxide gas were introduced into a pressure cell and adjusted to be above the equilibrium pressure of the hydrate and below the liquefaction pressure of the carbon dioxide at each temperature. After the formation of carbon dioxide hydrate, the residual gas was collected and the hydrate was recovered under the liquid nitrogen temperature to obtain the hydratebound gas. The carbon stable isotope ratios of carbon dioxide in the hydrate and residual gases were determined using a stable isotope mass spectrometer. Over a wide range from 243 K to 278 K, the δ^{13} C of residual gas was always larger than the hydrate phase, indicating that the carbon dioxide hydrate preferentially enclathrated lighter molecules. This trend is consistent with the results of a previous study. These results suggest that the equilibrium pressure of 13 CO₂ hydrate is slightly higher than that of ¹²CO₂ hydrate.

Keywords: gas hydrate, carbon dioxide, isotopic fractionation, equilibrium pressure, isotopologue

1. Introduction

Most naturally occurring gas hydrates are methane hydrates, but natural CO_2 hydrates also exist. The CO_2 hydrate sample was observed at the Okinawa Trough, where CO_2 seeps from the seafloor at a depth of 1335-1550 m and form crystals (Sakai et al., 1990). On the other hand, the possibility of CO_2 hydrates in the polar caps of Mars has long been discussed (Miller and Smythe, 1970). Natural gas hydrates retrieved at Lake Baikal also contain small amount of CO_2 . CO_2 is one of the main components of natural gas, however, little has been known about CO_2 as a guest gas of clathrate hydrate.

Since CO_2 is composed of carbon and oxygen, there are several isotopic species of CO_2 depending on the combination of stable isotopes (isotopologues: ¹²C, ¹³C, ¹⁶O, ¹⁷O, and ¹⁸O). For example, the abundance ratio of ¹³CO₂ is about 1.1% of the total CO_2 , and the rest is almost ¹²CO₂. In geochemistry, the stable isotope ratio of carbon is used as an indicator of the origin of CO_2 and related hydrocarbon gases. Luzi et al. (2011) first reported carbon stable isotope fractionation during the formation of CO_2 hydrate, and revealed that the CO_2 $\delta^{13}C$ in the hydrate phase was 0.9‰ lower than that in the gas phase at 268 K. This result indicates that $^{12}CO_2$ is more easily enclathrated in the hydrate phase than $^{13}CO_2$.

Stable isotope fractionation of the guest gas during the formation of gas hydrate is important information to discuss the formation, maintenance and decomposition processes of gas hydrate. In this study, CO_2 hydrate was synthesized and stable isotope fractionation was detected over a wide range of temperatures.

2. Material and methods

Research grade carbon dioxide gas (99.999%, Takachiho Chemical Industrial, but with a purity of about 99% as ${}^{12}CO_2$) was used as the guest gas. CO_2 hydrates were formed at 243 K, 255 K, 258 K, 263 K, 268 K, 273 K, 274 K, and 278 K. For the experiments below the ice point, 0.7 g of fine ice powder was filled in a pressure cell (volume: 30 mL) and set in a cold room at 255 K. For the experiment at 274 K, 5.3 g of distilled water was filled into another pressure cell (volume:

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145 mL) with a stirring device. After vacuuming the air in the cell at the temperature of liquid nitrogen, CO_2 was introduced and controlled above the equilibrium pressure of CO_2 hydrate and below the CO_2 liquefaction pressure at each temperature. After the formation of the crystals, residual gas was collected and then the crystals were collected at the temperature of liquid nitrogen. For the residual gas, the sampling tube was connected to the vacuum line and the pressure of the sample CO_2 was adjusted to the atmospheric pressure. The hydratebound gas was also obtained by dissociating the crystal in the vacuum line.

Each gas sample was introduced into a stable isotope mass spectrometer (Delta V, Thermo Fisher Scientific) by a syringe injection. The difference between the δ^{13} C of the hydrate-bound gas and that of the residual gas (δ^{13} C of the residual gas – δ^{13} C of the hydrate-bound gas, defined as $\Delta\delta^{13}$ C) was determined. A positive value of $\Delta\delta^{13}$ C indicates that light CO₂ (i.e. ¹²CO₂) prefers to be enclathrated.

3. Results

The CO₂ δ^{13} C of hydrate bound gas was always 1.0–1.5‰ lower than that of the residual gas in a wide range of temperatures from 243 K to 278 K. These results indicate that the hydrate phase tends to preferentially enclathrate ¹²CO₂. It also agrees with the previous work (Luzi et al., 2011) that reported 0.9‰ fractionation at 258 K. The temperature effect on carbon stable isotope fractionation tended to increase slightly with decreasing temperature, but this is not yet certain, considering the error of stable isotope measurements.

4. Discussion and conclusions

In the case of methane stable isotope fractionation during the formation of methane hydrate showed little carbon stable isotope fractionation of methane (Hachikubo et al., 2007). On the other hand, the hydrogen stable isotope fractionation of methane has been detected and methane δD in the hydrate-bound gas is $4.8 \pm 0.4\%$ smaller than that in the residual gas, suggesting that light CH_4 is concentrated in the hydrate phase, while CH_3D tends to remain in the gas phase.

These results can be explained by the difference in the equilibrium pressure of gas hydrate. The equilibrium pressure of CH_3D hydrate is slightly higher than that of CH_4 hydrate (Ozeki et al., 2018). It is easy to understand that methane is a mixture of isotopic species, even if they are the same methane. It is concluded that the difference in the equilibrium pressure of isotopologue gas hydrates causes the stable isotope fractionation of guest gases.

Applying the above discussion to the study of CO_2 hydrate, it is expected that the equilibrium pressure of ${}^{13}CO_2$ hydrate was slightly higher than that of ${}^{12}CO_2$ hydrate according to the results of carbon isotope fractionation.

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